

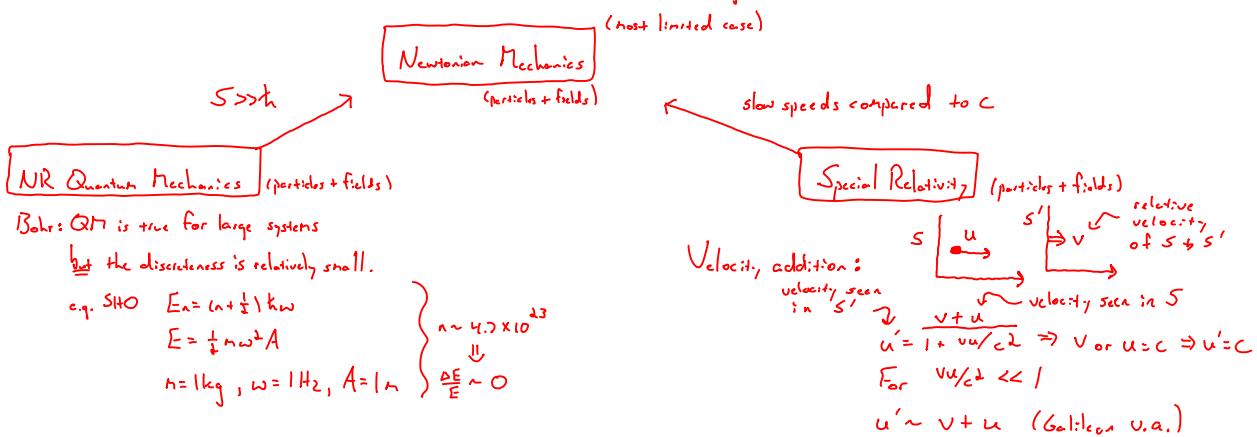
Why I like General Relativity (GR):

- GR is one of the few "brand new" subjects that students see late in the game (after they have a pretty solid understanding of basic physics).
- GR raises/questions some of the BIG questions in physics, e.g. cosmology, nature of spacetime.
- GR is 1/4 of the most perplexing issue confronting theoretical physics, i.e. Quantum Gravity.

What is "general" about GR and how is it related to Special Relativity (SR)?

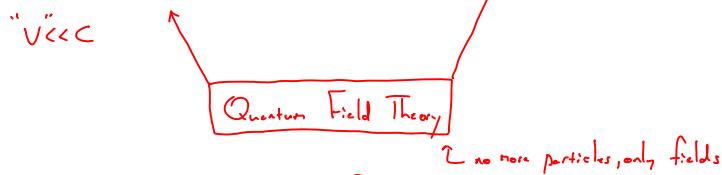
Many people think that GR is a generalization of SR ... and they are wrong! (but really, who could blame them?!)

To answer this we will first consider familiar correspondence principles (general \rightarrow limited).



Statistical: Classical behavior arises for many particles due to lack of wavefunction coherence (exceptions are condensates)

Path Integral:
Path weighted by $e^{\frac{i}{\hbar} \int L_{\text{free}} dt}$, so when $\hbar \ll S = \int L dt$
rapid fluctuations cancel except for extrema of S (Strudel), i.e. classical motion.



Okay, so where does GR fit in? All of the above (NH, QM, SR, QFT) are frameworks for mechanics. By themselves they do not constitute theories. To have a theory you need a framework plus degrees of freedom and their interactions.

Punchline: GR is a theory, not a framework. It is a theory of the gravitational interaction.

Two conclusions: a) There is no correspondence principle relating GR to SR.

b) GR is actually a generalization of Newtonian gravity (needed when $\frac{m}{R} \gtrsim \frac{c^2}{G}$).

Okay, but you still might ask how GR and SR are related. The answer is that SR is one particular solution of GR, i.e. a flat spacetime w/ no gravity acting.

In going beyond Newtonian gravity to GR we are going to be forced to reckon w/ a lot of new ideas, e.g. manifolds, curvature, tensors, etc.

Before it gets hairy though, we can paraphrase GR by analogy with another theory you know well,

Electrodynamics:

$$\begin{aligned} \vec{\nabla} \cdot \vec{E} &= \frac{\rho}{\epsilon_0} & \left. \right\} \text{Field Eqs.} \\ \vec{\nabla} \times \vec{B} &= \mu_0 \vec{J} + \mu_0 \epsilon_0 \frac{\partial \vec{E}}{\partial t} & \left. \right\} \text{Maxwell's Eqns.} \Rightarrow \text{Tells us how sources } (\rho, \vec{J}) \text{ create fields } (\vec{E}, \vec{B}) \\ \vec{\nabla} \times \vec{E} &= -\frac{\partial \vec{B}}{\partial t} & \left. \right\} \text{BI} \\ \vec{\nabla} \cdot \vec{B} &= 0 \end{aligned}$$

$$\begin{aligned} \vec{F}_{\text{ext}} &= q(\vec{E} + \vec{v} \times \vec{B}) \quad \text{Lorentz Force} & \left. \right\} \text{Tells us how particles react to the fields.} \\ \sum_i \vec{F}_i &= m \vec{a}_b & \text{N2L} \end{aligned}$$

This is an example of a background/test particle split which though not fundamental, is still very useful for computation and application!

General Relativity:

Einstein's Equation \rightarrow Tells us how sources create gravitational fields/curvature.

Geodesic Equation \rightarrow Tells us how a test particle responds to curvature.

So do we have a gravitational field? Actually no, not in the Newtonian sense. However, in describing the curvature of a spacetime we will need to introduce perhaps the single most important element of GR, the metric field.